

# Energy momentum tensor

Newtonian gravity	General Relativity
gravitational potential $\Phi(x^i, t)$	metric $g_{ab}$
Newton's law $\frac{d^2 x^i}{dt^2} = -\delta^{ij} \nabla_j \Phi$	Geodesic equation $\frac{d^2 x^a}{d\tau^2} = -\Gamma_{bc}^a \frac{dx^b}{d\tau} \frac{dx^c}{d\tau}$
Newtonian deviation $\frac{d^2 \xi^i}{dt^2} = -\delta^{ij} \left( \frac{\partial^2 \Phi}{\partial x^i \partial x^j} \right) \xi^k$	Geodesic deviation $(\nabla_u \nabla_u \xi)^a = -R^a{}_{bcd} u^b \xi^c u^d$
Tidal forces $\frac{\partial^2 \Phi}{\partial x^i \partial x^j}$	Riemann curvature $R^a{}_{bcd}$
Laplace's equation $\nabla^2 \Phi = 4\pi G \rho$	Einstein equations $G_{ab} = 8\pi G T_{ab}$

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# SR dynamics

In order to discuss how we can construct suitable stress-energy tensors, we need to understand the role of energy and momentum in relativity better.

We already know that, in absence of forces, a body moves according to (in the local inertial frame)

$$\frac{du^a}{d\tau} = 0$$

When there are forces present, we can introduce an analogue of Newton's 2<sup>nd</sup> law by writing

$$m \frac{du^a}{d\tau} = f^a$$

This can't be “derived”, but it satisfies key criteria;

- principle of relativity; it is in tensor form
- reduces to the force-free equation
- reduces to Newton's 2<sup>nd</sup> law at low velocities

# Four acceleration

It now makes sense to introduce the four acceleration as

$$a^a = \frac{du^a}{d\tau} \quad \Rightarrow \quad f^a = ma^a$$

In a general frame this leads to  $a^a = u^b \nabla_b u^a$ .

It is worth noting that the normalisation of the four velocity means that we must have

$$\frac{d(u^a u_a)}{d\tau} = 0 \quad \Rightarrow \quad u_a a^a = 0 \quad \text{that is} \quad u_a f^a = 0$$

In other words, there are only three independent equations of motion – just like in Newtonian physics.

# Four momentum

The equation of motion that we have written down leads naturally to the relativistic ideas of energy and momentum.

We define the four momentum as

$$p^a = mu^a \quad \Rightarrow \quad \frac{dp^a}{d\tau} = f^a$$

We also have

$$p^2 = p^a p_a = m^2$$

In a general inertial frame, moving with velocity  $v^i$  relative to the local inertial frame, we have

$$u^a = (\gamma, \gamma v^i) \quad \text{where} \quad \gamma = (1 - v^2)^{-1/2}$$

It then follows that, at low velocities, we have

$$p^t = \frac{m}{\sqrt{1 - v^2}} \approx m + \frac{1}{2}mv^2 \quad \text{and} \quad p^i \approx mv^i$$

Natural to interpret  $p^t = E$  as the energy and  $p^i$  as the three momentum.

# Energy

Solving our relations for the energy, we arrive at

$$E = \left(m^2 + p^2\right)^{1/2}$$

which shows that the rest mass is part of the energy of a relativistic particle.

For a particle at rest, and in the usual units, we have

$$E = mc^2$$

This is, perhaps, the most famous equation in all of physics...

Note: It may be more appropriate to refer to  $p^a$  as the “energy-momentum” four vector.

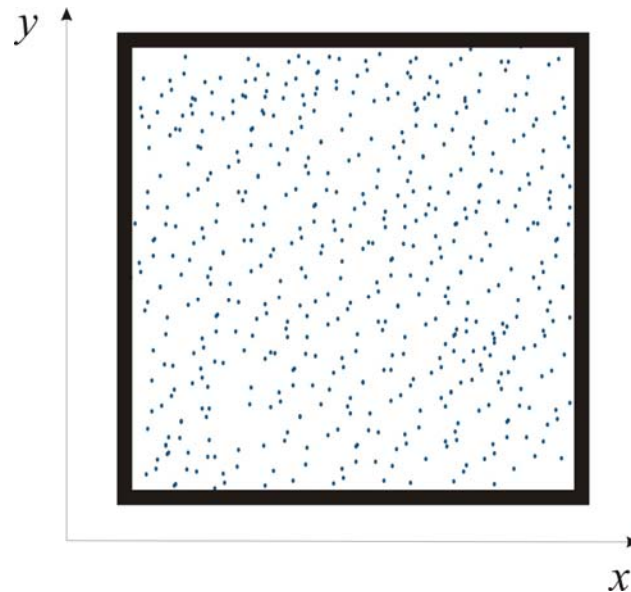
Also... it is the four momentum that is conserved in particle colliders like the LHC.

# Number density

In order to build our intuition of the description of matter in relativity, it is useful to start by considering the number density of a gas.

Let us consider a box containing  $N$  particles. At rest, the volume of the box is  $V_*$ . Then the number density of particles in the box is simply

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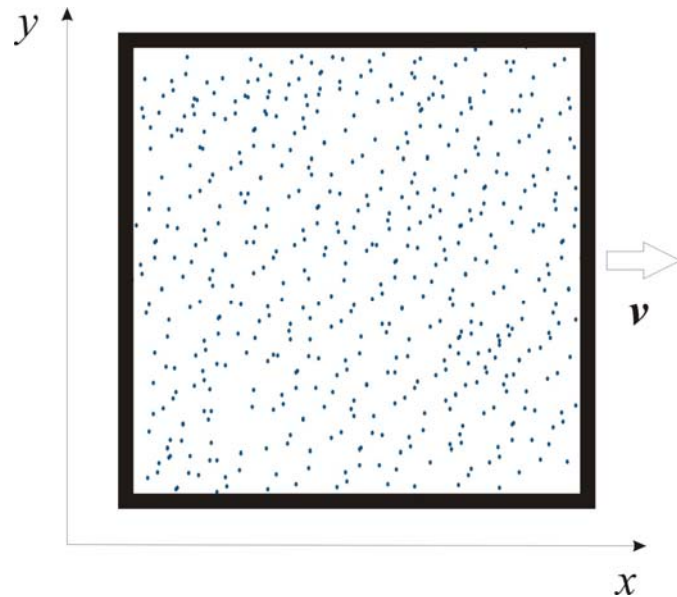
What happens if the box is moving?

Because of length contraction, the volume will be smaller;

$$V = (1 - v^2)^{1/2} V_*$$

but the total number of particles is the same, so the number density increases

$$\Rightarrow \frac{n}{\sqrt{1 - v^2}}$$



# Number conservation

We see that the number density is  $nu^t$ . This suggests that we should introduce the particle flux four vector as

$$n^a = nu^a$$

This means that we have

$$n^a = (n^0, n^j) = \left( \frac{n}{\sqrt{1-v^2}}, \frac{nv^i}{\sqrt{1-v^2}} \right)$$

Using the argument that leads to the conservation of particles in fluid dynamics (flux through surface of some volume...) we can show that

$$\partial_t n^0 + \nabla \cdot \mathbf{n} = 0 \quad \Rightarrow \quad \nabla_a n^a = 0$$

# Energy momentum tensor

We have seen how the number density current relates a scalar quantity with a volume. Let us now suppose that we want a similar argument for energy and momentum.

These are, however, given by the four-momentum. To relate this object with a volume, we need a tensor of rank 2.

This leads us to the energy-momentum tensor;

$$T^{ab} = \left( \begin{array}{c|c} \text{energy density} & \text{energy flux} \\ \hline \text{momentum density} & \text{stress tensor} \end{array} \right)$$

Consider the moving box, and assume that all particles are at rest with respect to the box. Then

$$\varepsilon = \text{energy density} = T^{00} = m n u^0 u^0 = m n \gamma^2$$

$$\pi^i = \text{momentum density} = T^{0i} = m n u^0 u^i = m n \gamma^2 v^i$$