

# 5. Boltzmann's Law

8<sup>th</sup> May

# Recap

- Lecture 0 - all common states of matter made of atom (KE v PE determines state)
- Lecture 1 - ideal gas  $PV=Nk_B T$ , from kinetic theory  $U = \frac{1}{2} n_d Nk_B T$
- Lecture 2 - zeroth law defines  $T$   
first law  $dU = \delta Q + \delta W$   
(for gases:  $dU = \delta Q - PdV$ )

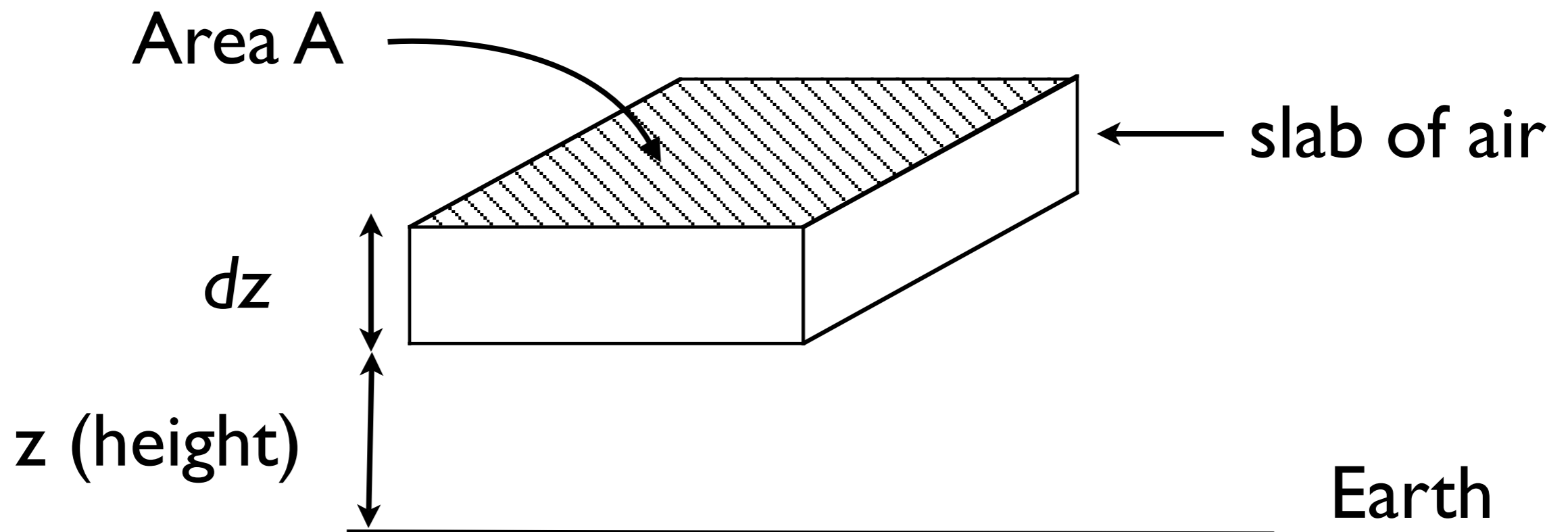
# Recap

- Lecture 3 - For gas  $dW = -PdV$  - path dependent  
e.g isothermal  $dW = -(Nk_B T) \ln(V_1/V_0)$ ,  
isobaric (followed by isochoric)  $dW = -(Nk_B T) (V_1/V_0 - 1)$
- Heat capacity (HC) define  $dQ = C dT$  - (also path dependent)  
(molar HC,  $dQ = N_m C_m dT$ , specific HC,  $dQ = m c dT$ )
- Ideal gas,  $C_v = n_d/2 Nk_B$  ( $C_{vm} = n_d/2 R$ )  $C_p = C_v + Nk_B$  ( $C_{pm} = C_{vm} + R$ )
- Adiabatic eqn of state  $PV^\gamma = \text{constant}$ , where  $\gamma = C_p / C_v$
- Lecture 4 - Phase Change - Large change in one state variable for a small change in another (usually indicating change in internal order and associated with latent heat).

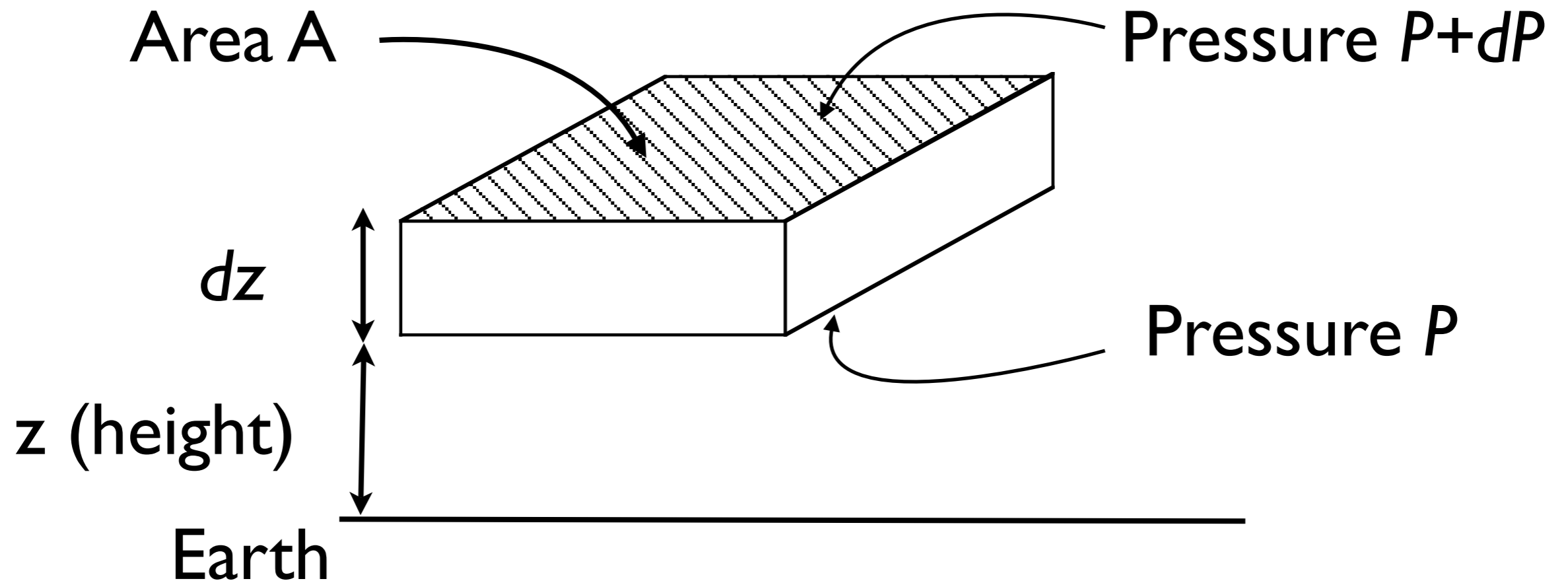
# 5.1 Isothermal Atmosphere

**Aim: To calculate air pressure as a function of height in equilibrium**

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**Figure 1**



①

②

③

$$\text{Net upward force} = PA - (P+dP)A - \rho A dz g$$

① pressure from below

② pressure from above (pressure is  $P+dP$  at  $z+dz$ )

[should find that  $dP < 0$ ]

③ weight of slab ( $\rho = \text{density}$ )

**but** in equilibrium, net force = 0

$$\rightarrow PA - PA - AdP - \rho A dz g = 0$$

$$\rightarrow \frac{dP}{dz} = -\rho g \quad (5.1.1)$$

$P$  falls with height.

Rewrite 5.1.1 using  $\rho(z) = n(z) m$

and assume isothermal ideal gas:  $P(z) = n(z) k_B T$

$$5.1.1 \quad \rightarrow \quad k_B T \frac{dn}{dz} = -mng$$

$$\rightarrow \frac{dn}{dz} = -\frac{mg}{k_B T} n$$

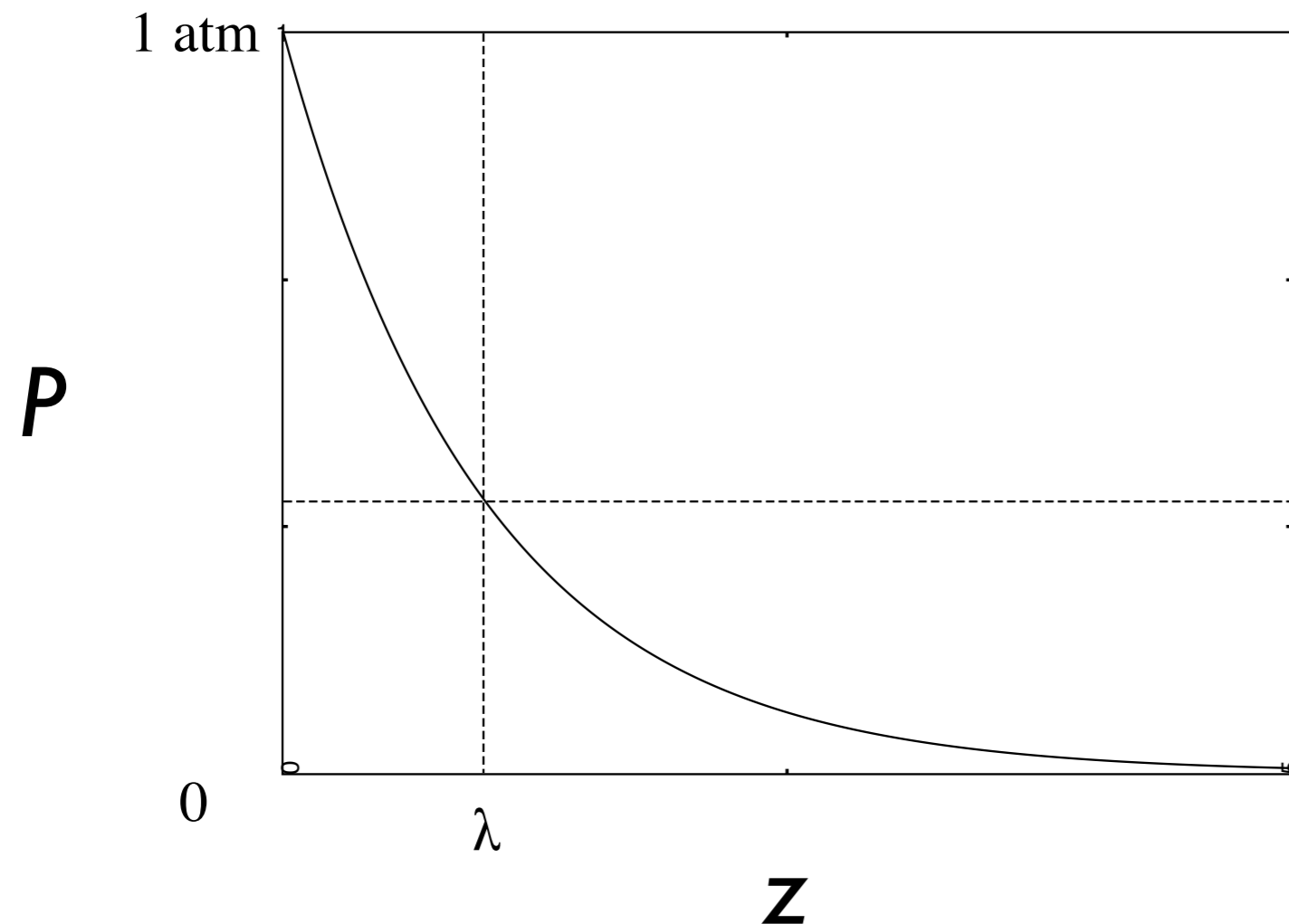


$$\therefore n = n_0 e^{-z/\lambda} \quad (5.1.2)$$

where  $n_0 = n(z = 0)$  [ie. at ground level],

$\lambda = k_B T / mg = \text{scale length}$

also  $P = nk_B T = P_0 e^{-z/\lambda}$

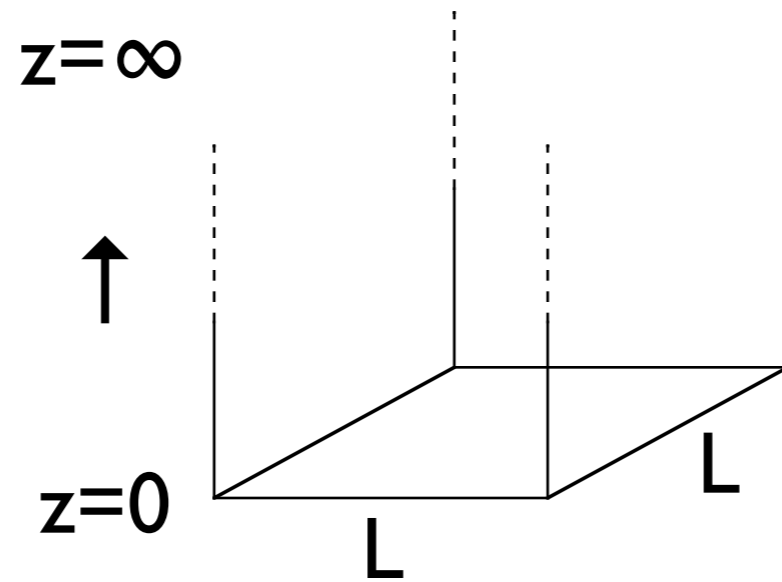


$\lambda = 8.5 \text{ km}$ , which is the height of Everest

human habitation only up to about  $5 \text{ km}$  (in Tibet & Andes)

So Everest ( $8.5 \text{ km}$ ) only conquered after lightweight oxygen canisters.

**Figure 2**



**Figure 3**

Consider a column of air, area  $L^2$ , infinitely high:

Number of particles between  $z$  and  $z+dz = n_0 \exp(-z/\lambda) L^2 dz$

$$\begin{aligned}
 \therefore N = \text{total number of particles} &= \int n_0 \exp(-z/\lambda) L^2 dz \\
 &= -n_0 L^2 \lambda [e^{-\infty} - e^0] \\
 &= n_0 L^2 \lambda \qquad (5.1.3)
 \end{aligned}$$

# Under pressure

$$n_0 = \frac{P}{k_B T} = \frac{10^5}{k_B 293} = 2.5 \times 10^{25} \text{ m}^{-3}$$

So total of particles above each  $\text{m}^2 = n_0 L^2 \lambda$

$$\approx 2.5 \times 10^{25} \times 8.5 \times 10^3 = 2.1 \times 10^{29}$$

The weight of these particles is  $1.66 \times 10^{-27} * 29 * 2.1 \times 10^{29} = 10^4 \text{ kg} = 10 \text{ tonnes!}$

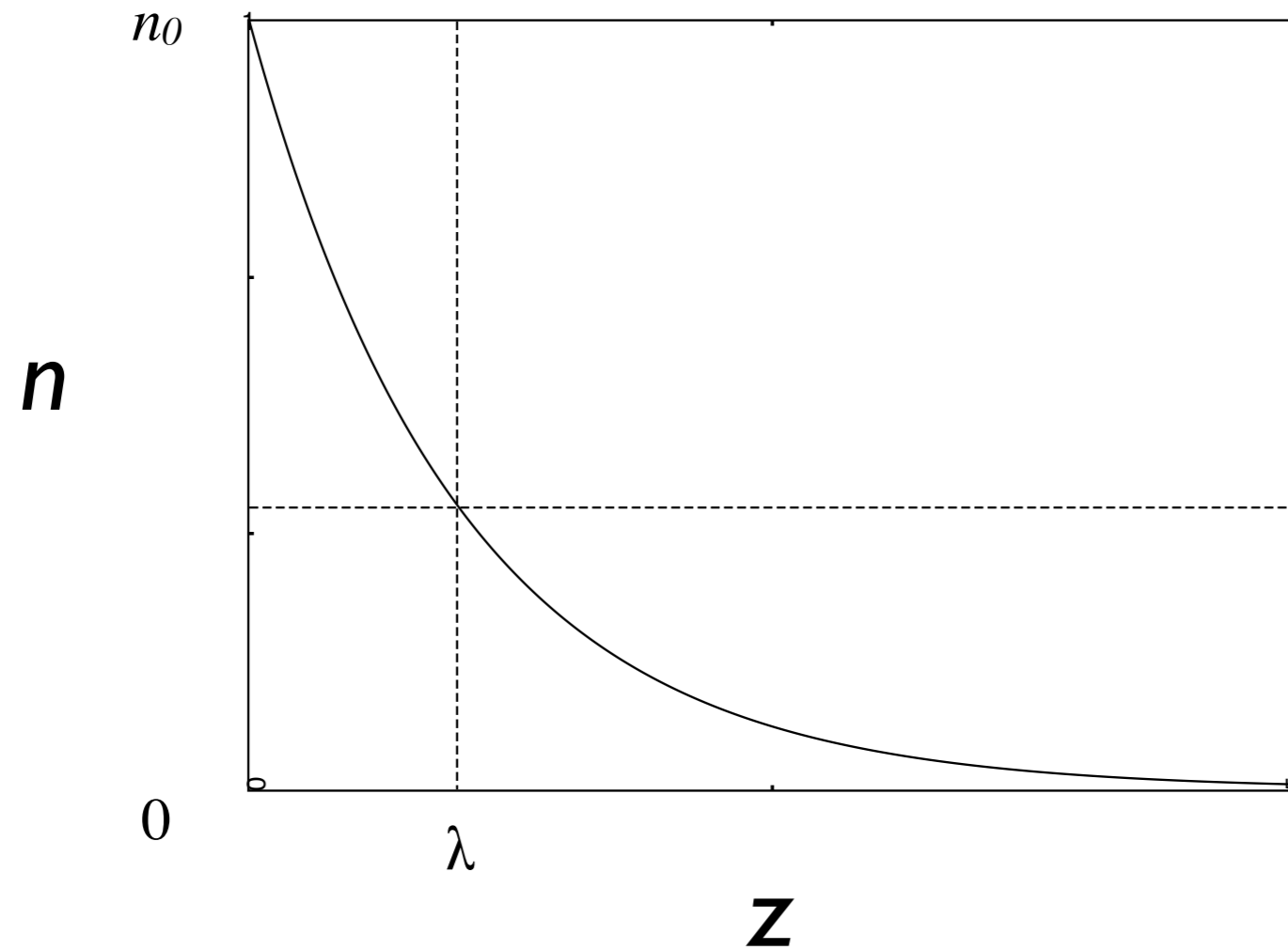
Of course the force due to  $10^4 \text{ kg} \approx 10^5 \text{ N/m}^{-2} =$   
atmospheric pressure.

Commercial airliners fly at about 32,000 ft 9-10km, which is why airliners are pressurised, but NOTE airliners only pressurised to 0.85 Atmospheres (to keep thickness required down)

# 5.2 Particle probabilities

$$n = n_0 e^{-z/\lambda}$$

(5.1.2)



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So Everest (8.5 km) only conquered after lightweight oxygen canisters.

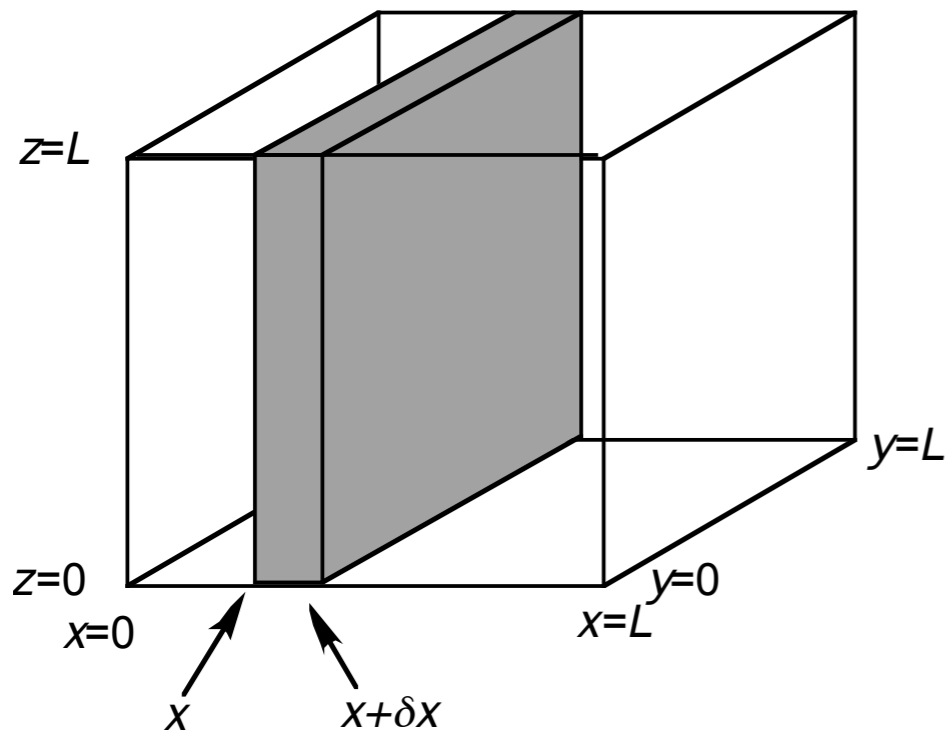
Probability of given molecule being in a region of the box = no. of particles in region /  $N$

e.g. if uniformly distributed  
 $n = N/L^3$

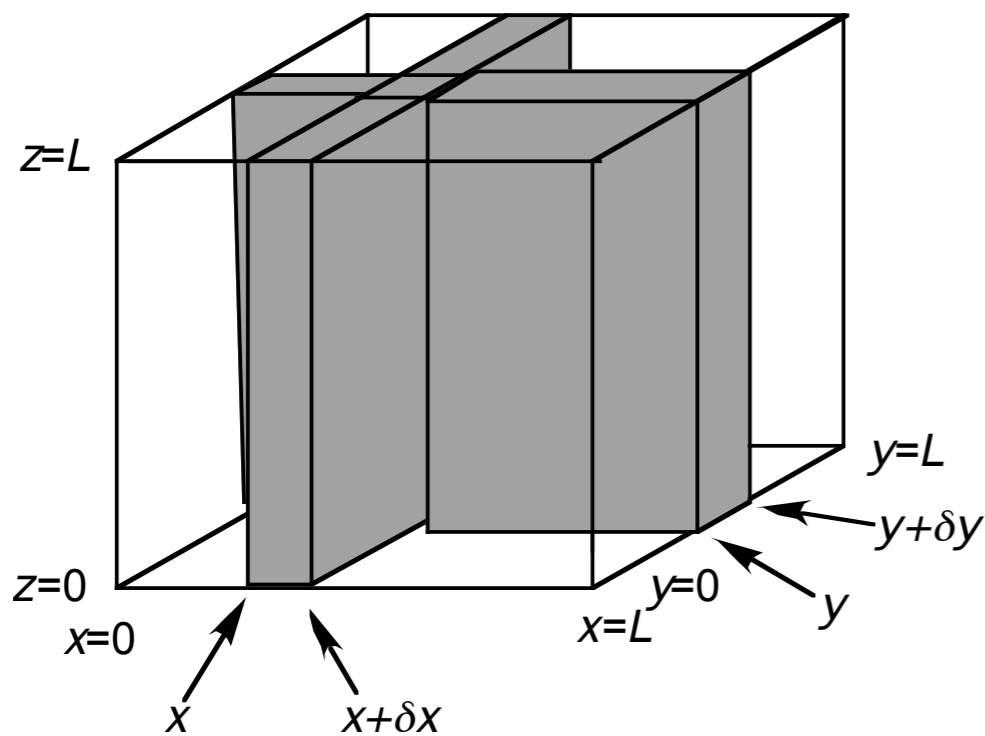
$\therefore$  number of particles between  $x$  and  $x + \delta x = n\delta x L^2$   
 $= (N/L^3)\delta x L^2$

so probability of given particles being between  $x$  and  $x + \delta x$ :

$$p(x)\delta x = \delta x/L$$



**Figure 4**



**Figure 5**

Number of such particles  
 $= N\delta x\delta y/L^2$   
 $= (N/L^3) \times L\delta x\delta y$   
 $(= n \times \text{volume of region})$

e.g. prob. of a particle being  
 between  $x$  and  $x + \delta x$  and also  
 between  $y$  and  $y + \delta y$

$$p(x, y) \delta x\delta y = \delta x\delta y/L^2$$

$\therefore$  if two events are independent,  
 prob. of both = prob(1)  $\times$  prob(2)

## For isothermal atmosphere:

5.2.1  $\rightarrow$  prob. of particles being between  $z$  &  $z + dz$  is  $p(z) dz$

$$\begin{aligned} p(z) &= (\text{no. of particles between } z \text{ \& } z + dz) / N \\ &= n_0 \exp(-z/\lambda) L^2 dz / n_0 L^2 \lambda \\ &= (mg/k_B T) \exp(-mgz/k_B T) dz \end{aligned} \quad (5.2.2)$$

**NB** Particles are actually moving through out the gas, so for any given particle, this is also the probability of where you would find it through-out its history.

Also note that equilibrium implies, stationary or constant state, but actually in a gas, though state variables are constant, underlying motion at atomic scale is far from stationary!



## 5.3 Boltzmann's Law

5.2.2  $\rightarrow$  prob of particles being between  $z$  &  $z + dz \propto \exp(-mgz/k_B T) dz$

but  $E(z) = mgz = \text{p.e.}$

so  $p(z) \propto p(E) \propto \exp(-E(z)/k_B T)$

IN FACT generally when in thermal equilibrium,

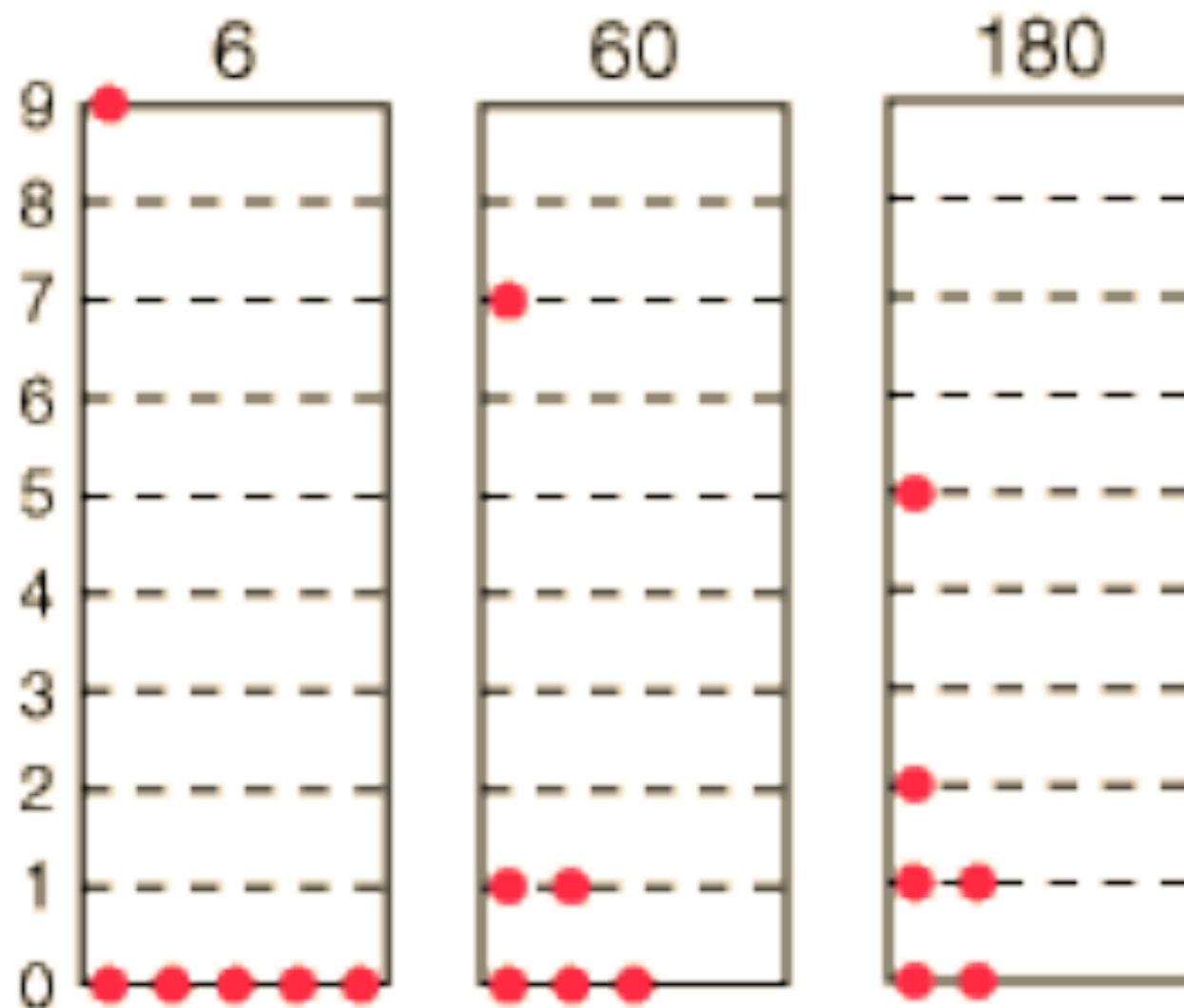
$p(E) \propto \exp(-E/k_B T)$  (Boltzmann's Law)

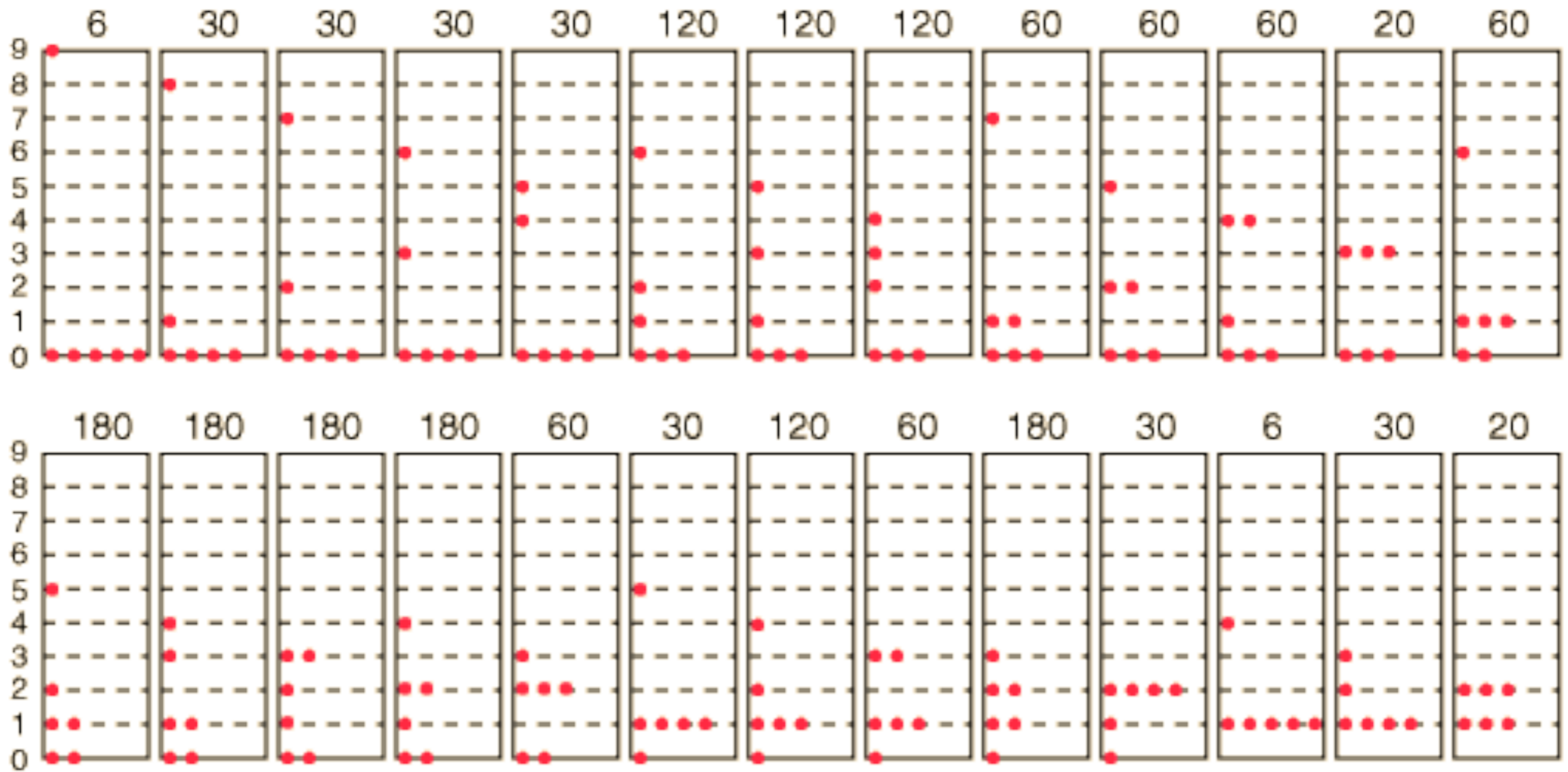
( $E$  can be any general potential)

# Counting states:

<http://hyperphysics.phy-astr.gsu.edu/hbase/quantum/disbol.html#c2>

**How many ways can you distribute 9 units of energy among 6 particles?**





**Answer: 26 distributions, but for distinguishable particles  
2002 possible arrangements**

Energy level number	Average
0	2.143
1	1.484
2	0.989
3	0.629
4	0.378
5	0.21
6	0.105
7	0.045
8	0.015
9	0.003

