Mechanics First Year Physics

Problem Sheet 9: Lectures 5.2 and 5.3

The following data are needed for some the questions on this problem sheet. $G = 6.67 \times 10^{-11}$ N m² kg⁻². Mass of the Earth = 5.98×10^{24} kg. Mass of the Sun = 1.99×10^{30} kg. Radius of the Earth = 6.37×10^{6} m. Mean radius of the Earth's orbit about the Sun = 1.49×10^{11} m. Mean radius of the Moon's orbit about the Earth = 3.84×10^{8} m.

Exercises

- 1. If the Earth was a perfectly smooth sphere (no mountains, trees, buildings, people, etc) without an atmosphere, an object could orbit just above the ground. Calculate the speed of such an object.
- 2. Show that a satellite in a circular orbit around a planet of mass M has an orbital period given by $T = \frac{2\pi r^{3/2}}{\sqrt{GM}}$, where r is the radius of the orbit.
- 3. Calculate the radius of a geostationary orbit around the Earth.
- 4. Uranus has a moon Oberon which has an orbital period of 1.16×10^6 s and a mean orbital radius of 5.84×10^8 m. Calculate the mass of Uranus.

Problems

- 5. (i) The Moon experiences gravitational forces from both the Earth and the Sun. Calculate the ratio of the magnitude of these two forces on the moon.
 - (ii) You should have found that the Sun exerts a larger force than the Earth. So, why isn't the Moon in orbit around the Sun?

- 6. (i) In Problem Sheet 6 (Q. 5) we considered the Bohr model of the Hydrogen atom, in which an electron follows a circular orbit around a proton. We found that the speed of the electron is given by $v = e/\sqrt{4\pi\epsilon_0 m_e r}$, where r is the radius of the orbit, and m_e is the electron mass. Given that the electrical potential energy of an electron distance r from a proton is $U = -\frac{e^2}{4\pi\epsilon_0 r}$, show that $|E| = |K| = \frac{1}{2}|U|$, where E and K are the electron's total mechanical energy and kinetic energy respectively. [We derived this relationship for the case of a circular gravitational orbit in Lecture 5.2. It applies in both cases because both the electrical force and gravity are inverse square ($\propto r^{-2}$) central forces.]
 - (ii) We found that the quantization of the electron's orbital angular momentum meant that only certain orbits were allowed, characterized by the principal quantum number, n. The radius of the n^{th} orbit is

$$r_n = n^2 \frac{\hbar^2 4\pi\epsilon_0}{m_e e^2}$$
 [see Problem Sheet 6, Q. 5.]

Calculate the total mechanical energy of an electron in the ground state (n = 1).

(iii) Hence find the ionization energy of Hydrogen (i.e., the energy required to remove the electron from the ground state to $r = \infty$). Give your answer in eV.

$$h = 1.05 \times 10^{-34} \text{ kg m}^2 \text{ s}^{-1}$$

$$\epsilon_0 = 8.85 \times 10^{-12} \text{ F m}^{-1}$$

$$m_e = 9.11 \times 10^{-31} \text{ kg}$$

$$e = 1.60 \times 10^{-19} \text{ C}$$

$$1 \text{ eV} = 1.60 \times 10^{-19} \text{ J}$$

- 7. In Lecture 5.2 we saw that energy conservation for an object of mass m in orbit around an object of mass M looks like the equation for an object executing 1-D (in the radial direction) motion in a "pseudo-potential" $U^* = \frac{L^2}{2mr^2} - \frac{GMm}{r}$.
 - (i) Show that the minimum of U^* occurs at a radius given by $r_c = \frac{L^2}{GMm^2}$ and that the minimum value is given by $U^*_{min} = -\frac{GMm}{2r_c}$.
 - (ii) Qualitatively describe the objects motion if its total mechanical energy is: (a) equal to U_{min}^* , (b) between U_{min}^* and zero, (c) greater than zero.

- 8. In Classwork I we found that the tidal force acting on a composite object consisting of two spheres (each of mass m and radius a) falling under gravity towards a black hole (of mass M) is given by $F_{tidal} = \frac{4GMma}{r^3}$, where r is the distance of the composite body from the black hole, and we have assumed r >> a. If the composite body was in orbit around the black hole it would still be constantly falling, so the expression for F_{tidal} would still apply. And, of course, any object will give rise to tidal forces, not just a black hole.
 - (i) Consider a composite object in orbit around a planet of mass M. Show that if the composite object is held together by gravity then it cannot exist below a minimum distance from the centre of the planet, given by $r_{min} = a(16M/m)^{1/3}$. (This minimum distance is called the Roche limit.)
 - (ii) Assuming that the planet and the composite body have equal densities, find the ratio of r_{min} to the radius of the planet.
 - (iii) Artificial satellites orbit the Earth well within the Roche limit. Why aren't they torn apart?

Numerical Answers

- 1. $7.91 \times 10^3 \ {\rm m \ s^{-1}}$
- 3. $4.22\times 10^7~{\rm m}$
- 4. $8.76\times 10^{25}~{\rm kg}$
- 5. (i) $F_{from Sun}/F_{from Earth} = 2.21$
- 6. (ii) -2.19×10^{-18} J, (iii) $+2.19 \times 10^{-18}$ J or 13.7 eV [Using more accurate values of the constants gives 13.6 eV]
- 8. (ii) 2.52.