

Question	Expected Answers	Further Guidance	Marks
1(a)(i)	Elliptical orbits (Sun at one focus) (Radius vector sweeps out) equal areas in equal times. $T^2 \propto R^3$ or equivalent		1 1 1
(ii)	Similarity- heliocentric Difference - Copernicus- circular orbits		1 1
(b)	$5.88^2/(3.55 \times 10^5)^3 = 1.12^2/R^3$ $R^3 = 1.623 \times 10^{15}$ $R = 1.18 \times 10^5$ km	Candidates <u>must</u> show their working	1 1 1
(c)	Perturbations (of gravitational field) Due to proximity of other moons		1 1 [10]
2.(a)(i)	(Measure of) brightness of star as seen from <u>Earth</u> . (Measure of) brightness of star as seen from (standard distance of) 10 pc.	Allow equations if terms explained.	1 1 1 1 1
(ii)	$m-M = 5 \lg (r/10)$ $r =$ distance in <u>parsecs</u>		1 1 1 1
(b) (i)	positions plotted correctly		1
(ii) 1	C		1
2	A		1
3	D		1
4	E		1 [11]
3.(a)	In an infinite And static universe sky would be bright at night because every line of sight ends on a star (constant rate of) expansion of universe Hubble law stated implies zero size in the finite past	Or other argument To 6 max	
(b)(i)	$1 \text{ pc} = 3 \times 10^{16} \text{ m}$ or $1 \text{ Mpc} = 3 \times 10^{22} \text{ m}$ $70 \text{ km s}^{-1} \text{ Mpc}^{-1} = 70 \times 1000 / (10^6 \times 3 \times 10^{16})$ $= 2.33 \times 10^{-18} \text{ s}^{-1}$	Or not static	6 1 1 1
(ii)	$T \approx 1/H_0 = 1/2.33 \times 10^{-18} = 4.3 \times 10^{17} \text{ s}$		1
(iii)	$s = vt = cT$ $= 3 \times 10^8 \times 4.3 \times 10^{17} = 1.3 \times 10^{26} \text{ m}$		1 1
(c)	uniform rate expansion	Or other valid assumption	1 [13]

	Marking Points	Additional guidance	Marks
4(a)	Speed of light constant (for all inertial observers) Laws of physics invariant for all inertial observers. In absolute space, c should vary with absolute velocity of observer, (but this is not the case - MM expt) mass increases with velocity as $v \rightarrow c$, $m \rightarrow \infty$, so infinite energy/force needed to reach c	Or other valid point Accept other valid arguments to 6 max Or $m = \gamma m_0$	6
(b) (i)	$t = s/v = 2000/(3 \times 10^8) = 6.6 \mu\text{s}$ comparison with $3 \times t_{1/2}$		1
(ii)	$1/2^3 = 1/8$ left $= 574/8 = 71.8$ in one hour	Allow 70 - 72	1
(iii)	muons travel close to speed of light so for an observer on the ground, the muons 'clock' will run much more slowly so its half life will be much greater.	Not just 'very fast'	1 1 1
(iv)	Observed fraction = $439/574 (= 0.765)$		1
(v)	correct substitution $\gamma = 7.92$		1
(vi)	$\ln f = 2000 \times \ln 2 / (0.992 \times 3 \times 10^8 \times 7.92)$ $= -0.268$ $\rightarrow f = 0.765$, (which agrees with expt)		1 1 1 [18]
5(a)	acceleration free obeying Newton's 1st law accelerating spaceship light moving perpendicular to acceleration follows curved path to accelerated observer on ship but straight path to inertial observer so by EP, light bends in gravity fields.	Allow valid alternatives.	6 max
(b)(i)	sky/Sun is very bright Only during an eclipse are stars near the sun visible.		1 1
(ii)	Gravity/curvature of spacetime strongest nearest Sun Diagram showing Light bent by Sun More bending for stars nearer Sun Apparent positions of stars shown	3 max	3
(c)	(i) Complete table correctly (ii) Plot points correctly Straight line <u>through origin</u> (iii) Straight line (through origin) indicates that α is proportional to $1/r$, QED (iv) $k = \text{gradient} = 1.8 \pm 0.15 R_0$ (v) $M_0 = 1.8/8.79 \times 10^{-31} = (2 \pm 0.1) \times 10^{30} \text{kg}$	correct calculation based on data only:2	1 1 1 1 1 1 [18]

			Marking Points	Additional guidance	Marks
6	(a)	(i)	Idea of zig-zag path 3-5 reflections, $i = r$, reflection at boundary Refractive index of light guide > air (allow denser) Provided $i \geq C$, TIR occurs.	Any 3	1 1 1
		(ii)	$\sin C = 1/n$ $C = \sin^{-1}(1/1.58) = 39.3^\circ$.		1 1
		(b)	(i)	$E = hc/\lambda = 6.63 \times 10^{-34} \times 3 \times 10^8 / 413 \times 10^{-9}$ $= 4.82 \times 10^{-19} / 1.6 \times 10^{-19} = 3.01 \text{ eV}$	Photon energy eV conversion
		(ii)	10^4 photons = $3.01 \times 10^4 \text{ eV}$ % conversion = $3.02 \times 10^4 \times 100\% / 1.5 \times 10^6 = 2\%$		1 1
	(c)	(i)	$hc/\lambda = 2.2 \times 1.6 \times 10^{-19}$ $\lambda = 6.63 \times 10^{-34} \times 3 \times 10^8 / (2.2 \times 1.6 \times 10^{-19})$ $= 566 \text{ nm}$	560- 570 nm	1 1
		(ii)	the work function is greater than the photon energy / no photoelectron emission	Or equivalent	1
		(iii)	$hf = \phi + \frac{1}{2}m_e v_{max}^2$ $hf - \phi = (3.02 - 2.2) \times 1.6 \times 10^{-19} = 1.31 \times 10^{-19} \text{ J}$ $v_{max} = \sqrt{(2 \times 1.31 \times 10^{-19} / 9.1 \times 10^{-31})} = (5.0 - 5.4) \times 10^5 \text{ ms}^{-1}$	Or $hf = \phi + \text{KE}$ Omit eV conversion: 2/3 KE = 3eV or 2.2eV: 1/3	1 1 1
	(d)	(i)	3^{12} $= 531000$	Allow $3^{13} =$ 1590000	1 1
		(ii)	$Q = ne = 531000 \times 1.6 \times 10^{-19} = 8.5 \times 10^{-14} \text{ C}$ $I = Q/t = 8.5 \times 10^{-14} / 3 \times 10^{-9}$ $= 2.8 \times 10^{-5} \text{ A}$	Q=e: 1/3	1 1 1